A Light-Scattering Study of Al₂O₃ Abrasives of Various Grit Sizes

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Abstract

We report light scattering phase function measurements for irregularly shaped Al₂O₃ abrasive powders of various grit sizes. Q-space analysis is applied to the angular scattering to reveal a forward scattering regime, Guinier regime, power law regime with quantifiable exponents, and an enhanced backscattering regime. The exponents of the power laws for Al₂O₃ abrasives decrease with increasing internal coupling parameter ρ' , which is in agreement with previous observations for other irregular particles. Unlike other dust particles previously studied showing single power laws under Q-space analysis, the largest three abrasives, for which $\rho' \geq 100$, showed a kink in the power law, which is possibly due to the higher degree of symmetry for the abrasives than for all the particles studied previously. Direct comparison of the 1200, 1000, and 800 grit abrasive scattering to scattering by corresponding spheres shows that the scatterings approximately coincide at the spherical particle $qR \approx \rho'$ crossover point. Furthermore, the scattering at the maximum $qR=2kR$ by the irregularly shaped abrasives is close to the geometric centers of the glories of the spheres.

22 1. Introduction

23 Light scattering from spherical particles has been well studied and understood [1,2];
24 however, there is no complete description and understanding of light scattering by irregularly
25 shaped particles. Much work including both experiment and theory has been conducted to study
26 irregularly shaped particles [3–18]. However, few studies have changed relevant parameters in a
27 systematic way which could lead to the uncovering of patterns. We chose to do light scattering
28 phase function measurements from irregularly shaped Al_2O_3 abrasives due to their selectable
29 sizes that allows for systematic variation of the size.

30 In this paper we describe experimental phase function measurements – scattered intensity
31 as a function of angle – of Al_2O_3 abrasives and analyze the data with Q-space analysis. In Q-
32 space analysis [19–21], the scattered intensity is plotted vs. the magnitude of the scattering wave
33 vector q , which has the unit of inverse length, or the dimensionless qR , where R is a radial
34 dimension, on a log-log scale . The magnitude of the scattering wave vector is derived as [22]

$$35 \quad q = 2k \sin \theta / 2 \quad (1)$$

36 where $k = 2\pi/\lambda$ with λ the wave length and θ is the scattering angle. Unlike the conventional
37 analysis (scattered intensity vs. θ) which yields nondescript, indistinguishable curves, Q-space
38 analysis uncovers a pattern of features common to all shapes. The common pattern consists of a
39 constant forward scattering lobe at small q , then with increasing q , a crossover, Guinier regime
40 near $q \approx R^{-1}$, a power law regime with quantifiable exponents, and finally an enhanced
41 backscattering regime [23]. The power law exponents decrease with increasing internal coupling
42 parameter ρ' . The internal coupling parameter ρ' [24] is defined as

$$43 \quad \rho' = 2kRm^2 - 1m^2 + 2 \quad (2)$$

44 where m is the refractive index.

45 We chose to study abrasive particles because they have irregular shapes, are readily
46 available and can be obtained in a range of sizes as specified by the manufacturer. Because the
47 abrasive action requires rolling between the two surfaces that are grinding, they have a roughly
48 spherical overall shape, as opposed to particles with large aspect ratios such as flakes or needles.
49 Thus we anticipate that scattering by abrasive particles might have some characteristics in
50 common with scattering by spheres. Our analysis below will directly compare the scattering by
51 the abrasives with scattering by spheres under the eye of Q-space analysis. Moreover, the
52 comparison will pay attention to both the forward and backward scattering properties in light of
53 the discussion above and the possible power law that might lie in between.

54

55 2. Experiment

56 A $\lambda/4$ wave plate, after a 532nm vertically polarized laser, with fast axis 45° from vertical
57 was used to produce circularly polarized incident light (the same as unpolarized light to the
58 detector). Our light scattering apparatus detects light scattered at 31 angles from 0.32° to 157°
59 simultaneously. The forward scattering detection contains 16 angles from 0.32° to 9.9° and is
60 based on a design by Ferri [25]. For the side scattering, a custom elliptical mirror (Optiforms)
61 collects light at 15 angles from 15.3° to 157° . The detail of our light scattering apparatus is
62 described in [26]. Similar as [26], a power meter was used to verify the experiment was under
63 the single scattering limit by following the method described in [27].

We used a dust generator to aerosolize the abrasives [26]. The dust generator had a three inch in diameter and three inch tall cylindrical chamber into which the dust was loaded. The bottom of the chamber had a stir bar which spun at 200 revolutions per minute. Simultaneously, oxygen blew in from the bottom of the chamber. The outgoing aerosolized dust particles were directed to the scattering volume by vinyl tubing. We measured the phase function of 6 different grits of white Al_2O_3 abrasive powders from Panadyne Abrasives. For each measurement, 10 grams of grit sample was loaded inside the dust generator chamber. Al_2O_3 is birefringent but the two refractive indices – 1.760 and 1.768 [28] – are very close to each other. Although they slightly change with wavelength, we treated the refractive index as 1.76 which would not significantly affect ρ' . The grit numbers and the average sizes in diameters of the abrasives were 1200 grit ($3\mu\text{m}$), 1000 grit ($4.5\mu\text{m}$), 800 grit ($6.5\mu\text{m}$), 600 grit ($9.3\mu\text{m}$), 400 grit ($17.3\mu\text{m}$), and 320 grit ($29.2\mu\text{m}$) as given by the manufacturer. The average sizes were determined by the openings of the screen through which the abrasives passed.

3. Al_2O_3 abrasives scattering results

We performed Q-space analysis on the angular scattering by the Al_2O_3 abrasives as shown from Figs. 1 to 6. The figures also contain corresponding optical microscope images. Each image has a 10 micron scale bar in the lower left and a scale bar in the upper right indicating the diameter given by the manufacturer for comparison to the image. The scattered intensity with arbitrary unit (a.u.) is plotted vs. $q \text{ cm}^{-1}$ on a double logarithmic plot; this is Q-space analysis. For the smaller sizes, grits 1200, 1000, and 800 (note that the larger the grit number, the finer the powder), Q-space analysis reveals an indication of a constant forward scattering lobe at smallest

q followed by, with increasing q, a Guinier regime, a power law regime and an enhanced backscattering regime. The power law exponents are labeled in each graph. For the larger sizes, grits 600, 400 and 320, the forward scattering and Guinier regimes could not be seen since the large size requires angles smaller than our minimum angle of 0.32°. Also, for the larger grits, a new feature appears starting perhaps with 800 grit and then growing with increasing size: a kink near $q \approx 2 \times 10^4 \text{ cm}^{-1}$. This kink perturbs the power law functionality.

We have verified by calibrating the forward and side detectors to each other that the kink is not due to the misalignment or malfunction of our setup. This is substantiated by the fact that the kink does not exist for small abrasives and starts to develop with increasing sizes. Thus we are confident that the kink is an experimental fact.

Figure 1. Q-space analysis of scattering by 1200 grit Al_2O_3 abrasives and the corresponding image under an optical compound microscope.

99

100 **Figure 2.** Q-space analysis of scattering by 1000 grit Al_2O_3 abrasives and the corresponding
 101 image under an optical compound microscope.

102

103 **Figure 3.** Q-space analysis of scattering by 800 grit Al_2O_3 abrasives and the corresponding image
 104 under an optical compound microscope.

105

106 **Figure 4.** Q-space analysis of scattering by 600 grit Al_2O_3 abrasives and the corresponding image
 107 under an optical compound microscope.

108

109 **Figure 5.** Q-space analysis of scattering by 400 grit Al_2O_3 abrasives and the corresponding image
 110 under an optical compound microscope.

111

112 **Figure 6.** Q-space analysis of scattering by 320 grit Al_2O_3 abrasives and the corresponding
 113 image under an optical compound microscope.

114 As stated above, the size labeled in each figure is the average size of each abrasive
 115 sample given by the manufacture. However, light scattering tends to see the bigger particles of an
 116 ensemble of particles because bigger particles scattering more light. Hence, we need to
 117 determine the sizes from light scattering. Since the abrasives have irregular shapes, it is hard to
 118 define the sizes. However Guinier analysis of light scattering can yield the radius of gyration R_g
 119 which is exactly definable for any shape as

120
$$R_g^2 = \frac{1}{M} \int_0^M r^2 \rho(r) dr \quad (3)$$

121 where $\rho(r)$ is the mass density at position r of an abrasive particle. In the limit $\rho' \rightarrow 0$ and when
 122 $qR_g \ll 1$, the Guinier equation [29] describes the forward scattering as

123
$$I(q) = I(0) \left[1 - \frac{1}{3} q^2 R_g^2 \right] \quad (4)$$

124 When $\rho' > 1$, the Guinier equation form still holds but the radius of gyration inferred from Eq. (4)
125 is not the true radius of gyration [30]. For spheres when $\rho' > 30$, the ratio of the true to measured
126 radii of gyrations is about 0.85 ($\pm 10\%$). Since a study similar to [30] has not been performed for
127 non-spherical particles, and since the abrasive powders have a roughly spherical shape in some
128 manner, we will apply this correction to the R_g we determine from a Guinier analysis of our
129 abrasive light scattering.

130 The Guinier analysis [29] to determine the light scattering radius of gyration is shown in
131 Fig. 7. Note the paucity of data in the Guinier regime which warns us not to expect great
132 accuracy. We could only apply Guinier analysis to the 1200, 1000 and 800 grits which have
133 relatively small sizes to ensure $qR_g < 1$. The 1200 grit has several points for the Guinier analysis.
134 From the slope the radius of gyration of 1200 grit is $2.4\mu\text{m}$. For 1000 and 800 grits, the data are
135 more limited. We put two boundary slopes for each set of data. The fit lines (solid lines) with
136 smaller slopes only include two data points when $qR_g < 1$, where the Guinier analysis is, in
137 principle, applicable; the fit lines (dashed lines) with larger slopes are the fits for all three data
138 points in each data set. According to the fit line slopes, the ranges of the radii of gyration are
139 between $3.7\mu\text{m}$ and $3.9\mu\text{m}$ for grit 1000 and between $6.4\mu\text{m}$ and $6.9\mu\text{m}$ for grit 800. From
140 Guinier analysis, the measured radii of gyration of the 1000 and 800 grits were $3.8 \pm 0.1\mu\text{m}$ and
141 $6.6 \pm 0.3\mu\text{m}$. Given the uncertainties of radii of gyration of 1000 and 800 grits, the radius of the
142 gyration of grit 1200 is $2.4 \pm 0.1\mu\text{m}$. After the factor of 0.85 correction, the true radii of gyration
143 were $2.0\mu\text{m}$, $3.2\mu\text{m}$, and $5.6\mu\text{m}$, respectively, without considering the uncertainties.
144 Mathematically, R_g is the root mean square radius which is smaller than the visible perimeter.
145 The exact relation between R_g and the visible perimeter depends on the particle shape. Since we

do not know the exact shape, we again make an approximation based on spheres where $R_g = 35R$. Therefore, the optical perimeter radii of the 1200, 1000, and 800 grits were $2.6\mu\text{m}$, $4.2\mu\text{m}$, and $7.2\mu\text{m}$, respectively; the optical perimeter diameters were $5.2\mu\text{m}$, $8.4\mu\text{m}$, and $14.4\mu\text{m}$, respectively. They are bigger than the sizes labeled in Figs. 1 – 3. Again, we have to consider the fact that light scattering tends to see the bigger particles of an ensemble of particles. The manufacturer claimed the average sizes of 1200, 1000, and 800 grits were $3\mu\text{m}$, $4.5\mu\text{m}$, and $6.5\mu\text{m}$ correspondingly. However, the light scattering determined particle sizes were $5.2\mu\text{m}$, $8.4\mu\text{m}$, and $14.4\mu\text{m}$ correspondingly which were twice as big as the sizes given by the manufacturer. Therefore, we made an assumption that the light scattering determined sizes for grit 600, 400, and 320 would be twice as big as the sizes given by the manufacturer as well. Unfortunately, we could not determine the light scattering determined sizes due to the request of smaller angle detection for grit 600, 400, and 320. When we calculated ρ 's for the abrasives, we used the sizes twice as big as the ones given by the manufacturer.

Figure 7. Guinier analysis on Al_2O_3 abrasives of 1200, 1000, and 800 grit. The 1200 grit has several points for the Guinier analysis which indicates the radius of gyration is $2.4\mu\text{m}$. The data for 1000 and 800 grits are limited. The solid fit lines with smaller slopes only include two data points when $qR_g < 1$; the dashed fit lines with larger slopes are the fits for all three data points in each data set. According to the fit line slopes, the ranges of the radii of gyration are between $3.7\mu\text{m}$ and $3.9\mu\text{m}$ for grit 1000 and between $6.4\mu\text{m}$ and $6.9\mu\text{m}$ for grit 800.

4. Discussion

4.1 Power law exponent

In our recent, comprehensive work to study light scattering by irregularly shaped particles [23], we found that every particle type studied showed a uniform power law. Now, to the contrary, the three largest grits studied here show a kink in the power law. We will address this kink below when we compare abrasive scattering to spherical particle scattering. That previous work also showed that the power law exponents had a quasi-universal behavior with respect to the internal coupling parameter ρ' . In Fig. 8 we include the exponents for the grits, notwithstanding the kink, along with all the other exponents determined in that work (Fig. 8 is a reproduction of Fig. 9 from [23] with the six grits included). This demonstrates that the power law exponents from the grits follow the overall trend discovered in [23]. Note that the three largest grits, which showed the kink, are in a regime of $\rho' \gtrsim 100$ which was not explored in our previous work, save for the point at $\rho' \simeq 730$, the data for which did display a uniform power law.

180

181 **Figure 8.** Empirical correlation of Q-space power law exponents with the internal coupling
 182 parameter ρ' for a wide range of irregularly shaped particles [23]. The exponents for the Al_2O_3
 183 abrasives shown as open triangles fit well to the quasi-universal behavior.

184

185 4.2 Comparison between spheres and Al_2O_3 abrasives

186 In this section we make a comparison between the scattering from the irregularly shaped
 187 abrasives and perfect spheres with a polydisperse size distribution, using the perceptive eye of Q-
 188 space analysis. The comparison is semi-quantitative at best, but we believe it has value and
 189 points the way for future research. We compare grits and spheres of the same size and refractive

190 index, hence with the same ρ' . Only the three smallest grits have well defined forward scattering
 191 and Guinier regimes to match to the same regimes for the spherical particles. To achieve the
 192 match we generated Mie solutions using Philip Laven MiePlot [31] to find a sphere radius that fit
 193 best to the abrasives Guinier regime for each grit. This was done under the condition of a
 194 spherical particle size distribution width of 1.2, which is used to eliminate the Mie scattering
 195 ripples.

196 Figure 9 (a) – (c) show the comparisons between the spheres shown as dashed lines and
 197 the abrasives data shown as solid circles, where the scattered intensity (normalized at $qR \ll 1$) is
 198 plotted vs. qR . In Fig. 9 the abrasives scattering and the Mie scattering are forced to fit to each
 199 other in the Guinier regime so that we could see how they compare at large qR .

Figure 9. Comparisons between scattering by spheres and the 1200, 1000, and 800 grit abrasives. The scattered intensity (normalized at $qR \ll 1$) is plotted vs. qR for each grit. Spheres and abrasives were forced to fit in the Guinier, $qR \ll 1$, range. Also indicated are the $qR \approx \rho'$ crossover points for the sphere scattering.

The Q-space functionalities for spheres is as follows [24]: when $\rho' < 1$ (the Rayleigh-Debye-Gans (RDG) limit) a qR^{-4} functionality follows the Guinier regime. With increasing ρ' , a qR^{-2} functionality appears after Guinier regime and then this falls back to a qR^{-4} functionality when $qR \gtrsim \rho'$. As ρ' becomes larger, the qR^{-2} functionality breaks into a qR^{-3} regime followed by a relatively flat regime which then crosses over to a qR^{-4} functionality when $qR \gtrsim \rho'$. We will call this the “ $qR \approx \rho'$ crossover”. These functionalities are apparent in Fig. 9.

The comparison of the abrasive scattering to the spherical particle scattering in Fig. 9 uncovers some interesting features. At the $qR \approx \rho'$ crossover, the scattering of the abrasives roughly touches the scattering due to the spheres. The end points of the scattering, i.e. the scattering at the maximum $qR = 2kR$, by the irregularly shaped abrasives are close to the geometric centers of the glories of the spheres. Along with these features, the rough $(qR)^{-3}$

immediately after the Guinier regime and the qR -4 functionality for $qR \gtrsim \rho'$ are lost. One could say that the “roughening” of the spherical shape into that of the abrasives smoothed the intensity versus qR plot to leave a Guinier hump plus a single power law that is anchored by the Guinier regime and the $qR \simeq \rho'$ crossover. The same comparison was made in [32] between Gaussian Random Spheres (GRSs) and spheres to find the same result for the ending points, especially for large ρ' . Unfortunately, we cannot make direct comparison between GRSs and our abrasives because the smallest ρ' for our Al_2O_3 abrasives is larger than any GRS that we can currently solve for using computational techniques.

For the other three abrasives, 600, 400, and 320 grits, because the Guinier regime was not in the detectable range, we could not make direct comparisons to spheres. For spheres, the qR -3 regime is a consequence of the crossover of the three dimensional spherical particle scattering to two-dimensional, Fraunhofer diffraction [33] which must occur when the sphere becomes very large and refractive. We expect a similar 3d to 2d crossover would occur for any shape at large enough ρ' . However, the regime before the kink for the abrasives does not show a qR -3 functionality. The kink that occurs in the scattering for these three grits developed with increasing size, thus it must be related to size. Note that Fig. 8 shows that these three grits are in a regime of large ρ' that was not covered in the previous work (save for the point at $\rho'=730$ which was definitely a power law). Obviously this large ρ' region needs further study. If large ρ' is not the cause of the kink, we speculate that it could be due to the higher degree of symmetry for the abrasives than for all the particles studied previously.

237 5. Conclusions

238 We have measured the scattered light intensity for six different sizes of Al_2O_3 grits over a
239 wide range of scattering angles starting at 0.32° . Q-space analysis of the scattering reveals a
240 forward scattering regime at small q , followed by a crossover, Guinier regime near $q \approx R^{-1}$, a
241 power law regime with quantifiable exponents, and an enhanced backscattering regime. The
242 power law exponents decrease with increasing internal coupling parameter ρ' , which fits well to
243 our previously observed quasi-universal behavior of the power law exponents with respect to ρ'
244 for irregularly shaped particles. However, unlike that previous work, the largest three abrasive,
245 for which $\rho' \geq 100$, showed a kink in the power law. We directly compared the scattering of
246 1200, 1000, and 800 grit abrasives with corresponding spheres by finding the best fit in Guinier
247 regime. At large qR , the scattering of the abrasives approximately coincides with that of the
248 spheres at the $qR \approx \rho'$ crossover point of the scattering from spheres, and the abrasive scattering
249 follows a power law after the Guinier regime whereas the spherical particle scattering is much
250 more complex. The scattering at the maximum $qR = 2kR$ by the irregularly shaped abrasives is
251 close to the geometric centers of the glories of the spheres.

252 Finally, it is interesting to ponder why Q-space analysis uncovers features not seen with
253 conventional plotting versus the scattering angle θ . The most important feature is that q has units
254 of inverse length, and thus inverse q has units of length; it is the resolution length scale of the
255 scattering experiment. Thus at small enough q , the resolution scale q^{-1} is larger than the size of
256 the particle and so the scattering cannot resolve the particle thus shape is not a factor. Given this
257 perspective, one can realize that a change in q will not change the scattering so long as q^{-1} is
258 greater than the size. Continuing to larger q , as q^{-1} passes through the particle size, functionality

259 must begin and that is the Guinier regime. Thereafter, the particle is well resolved by the
260 scattering as a two dimensional object, regardless of shape, with two dimensional diffraction
261 which has a q^{-3} Porod regime. This was approximately, but not completely obtained, for the
262 particles here. At yet larger q , the scattering resolution becomes sharp enough to see the true
263 three dimensional nature of the particle which yields a q^{-4} Porod regime. Finally, at the largest q ,
264 enhanced backscattering occurs even for non-spherical particles for reasons we do not know.
265 These concepts for Q-space analysis can be found in more detail in reference [34].

266

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